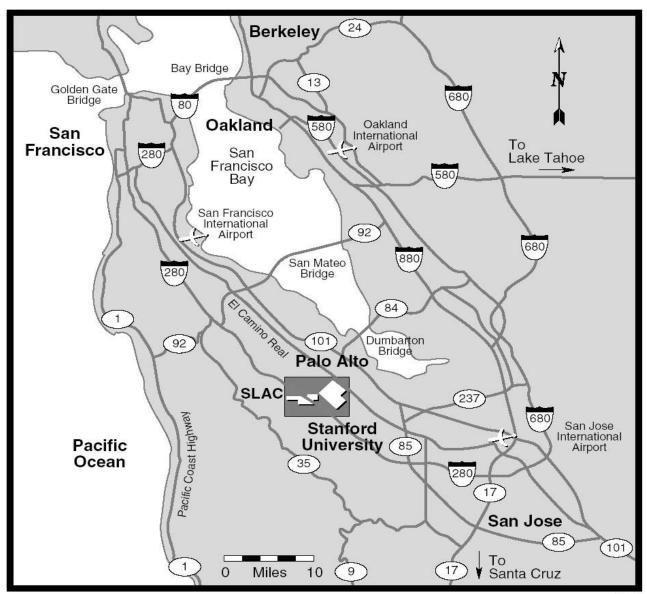
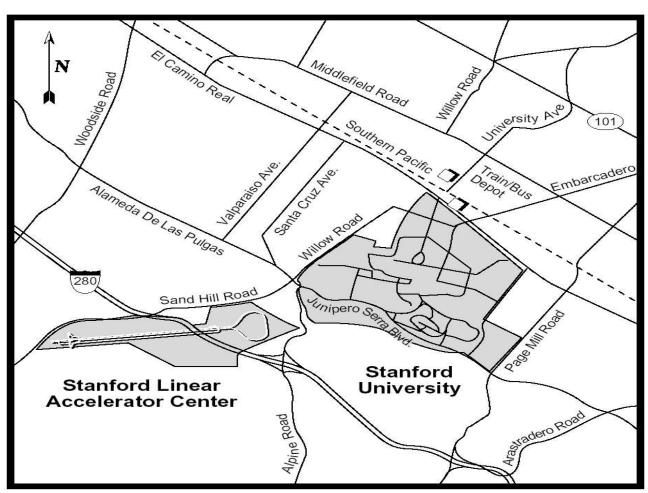


©1994 Magellan Geographix Santa Barbara, CA (800) 929-4MAP

Stanford Linear 4 Center



SLAC



Stanford Area Map



SLAC – основные этапы развития

- 1962 Начало сооружения ускорителя
- 1966 Начало физических экспериментов
- 1968 Получение первых свидетельств существования кварков
- 1972 Начало экспериментов на установке SPEAR
- 1973 Начало работы с источником синхротронного излучения
- **■** 1974 Открытие ψ- частицы
- 1976 Открытие с-кварка и т лептона
- 1976 Присуждение Нобелевской премии B.Richter за открытие J/ψ
- 1980 Начало физических экспериментов на установке PEP
- 1989 Начало физических экспериментов на установке SLC/SLD
- 1990 Присуждение Нобелевской премии R.Taylor за открытие кварков и структуры нуклонов
- 1995 Присуждение Нобелевской премии M.L.Perl за открытие τ лептона

SLAC – базовые установки.

- Линейный ускоритель электронов и позитронов. Длина 3 км.
 Максимальная энергия пучка 50 GeV.
- SPEAR накопительное кольцо частиц. Диаметр 80 м. Использовалось для организации встречных пучков е⁺е⁻с энергией 8 GeV. Сейчас используется как источник синхротронного излучения с энергией 3 GeV.
- PEP Накопительное кольцо. Диаметр 800 м. Использовался для организации встречных пучков е⁺е⁻ с энергией 30 GeV. Переделан в «ассиметричную В-фабрику (Е_{см}=10 GeV)» PEP-II/BaBar детектор.
- Линейные встречные пучки SLC и универсальный детектор SLD E_{см}=100 GeV.

SLAC

Facts at a Glance (as of 11/00)

- Budget 184 million \$
- Staff (FTE) 1314
- Users 2904
- User Institutions

Universities 147

Industry 46

Government Labs 30

Foreign162

1962

SLAC Ground Breaking

1966

Linac Begins Operation



1968

Quarks Discovered in Nucleon Nobel Prize 1990

1974-76

Charm Discovered at SPEAR Nobel Prize 1976





Tau Lepton Discovered at SPEAR Wolf Prize 1983; Nobel Prize 1995



B Meson Lifetime Measured at PEP



Limit of Three Quark Generations Measured at SLC

1991

Operation with Polarized Zs at SLC/SLD

1994

Construction Begins on (PEP-II) B-Factory

м

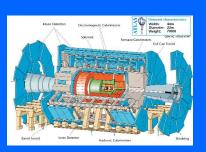
Как нашли структуру в адроне (кварки)?

Идеология таких измерений восходит к опытам
 Резерфорда в которых обнаружили структуру атома.
 Для таких измерений прежде всего необходимо найти достаточно «тонкий» щуп – пробник.

Зачем нужны ускорители - 1?

1. Длина волны "луча" с помощью которого исследуется "объект" должна быть меньше размера "объекта"!





$\lambda << h/P = hc/E$

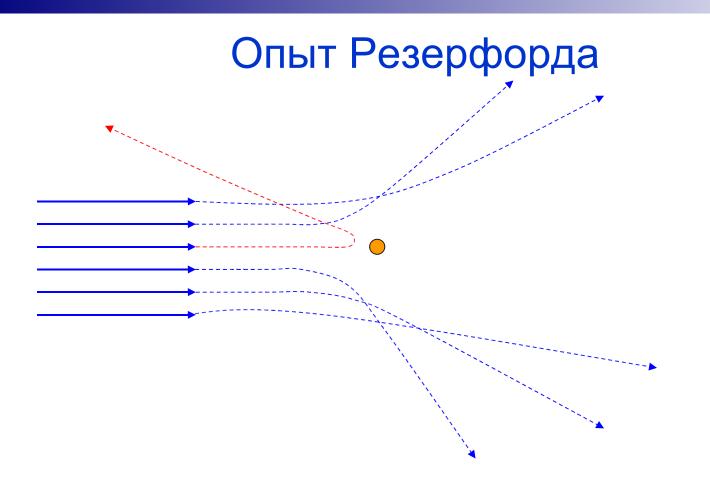
Объект	Размер	Энергия "луча"
Атом	10 ⁻¹⁰ м	10-5 ГэВ (10 эВ)
Ядро	10 ⁻¹⁴ м	10 ⁻² ГэВ (10 МэВ)
Нуклон	10 ⁻¹⁵ м	0.1 ГэВ
Кварк	??	> 1 ГэВ



м

Как нашли структуру в адроне (кварки)?

- Одно из преимуществ линейного ускорителя он может работать начиная с малой общей длины (энергии), которая будет возрастать по мере увеличения длины ускорителя
- Решающее преимущество электронного ускорителя в том, что электрон «точечная частица».



1990 Nobel Prize in Physics

The prize was awarded jointly to:

- •Friedman, Jerome I., U.S.A., Massachusetts Institute of Technology, Cambridge, MA, and
- •Kendall, Henry W., U.S.A., Massachusetts Institute of Technology, Cambridge, MA, and
- •Taylor, Richard E., Canada, Stanford Linear Accelerator Center, Stanford University, Stanford, CA.

"for their pioneering investigations concerning deep inelastic scattering of electrons on protons and bound neutrons, which have been of essential importance for the development of the quark model in particle physics

Обнаружение τ- лептона

Matter

$\left[\begin{array}{c} \mathbf{u} \\ \mathbf{d} \end{array}\right] \left[\begin{array}{c} \mathbf{c} \\ \mathbf{s} \end{array}\right] \left[\begin{array}{c} \mathbf{t} \\ \mathbf{b} \end{array}\right]$ Кварки

 $\left|\begin{array}{c|c}e\\v_{0}\end{array}\right|\left|\begin{array}{c|c}\mu\\v_{0}\end{array}\right|\left|\begin{array}{c}\tau\\v_{0}\end{array}\right|$ Лептоны

Interactions

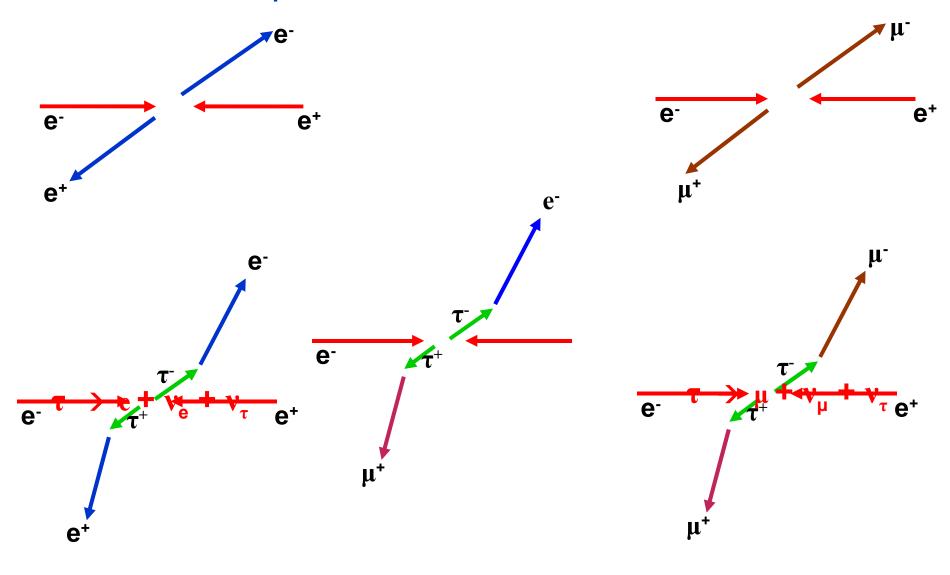
e – stable;
$$m_e \sim 0.5 \text{ MeV}$$

$$\mu \rightarrow e + \nu_e + \nu_\mu \qquad m_\mu \sim 100 \text{ MeV}$$

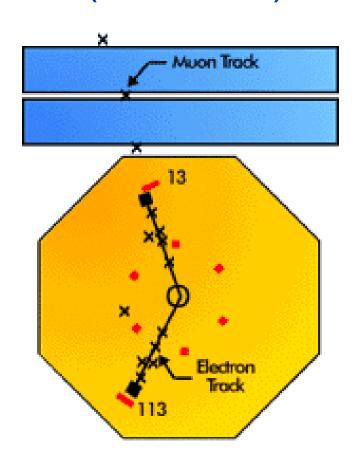
$$\tau \rightarrow ?e + \nu_e + \nu_\tau \qquad m_\tau > m_\mu$$

$$\tau \rightarrow \mu + \nu_\mu + \nu_\tau$$

Стратегия поиска au- лептона



Установка для поиска τ- лептона (MARK-III)



1995 Nobel Prize in Physics

The prize was awarded jointly to:

•Perl, Martin L., U.S.A., *Stanford Linear Accelerator Center*, Stanford University, Stanford, CA

"for pioneering experimental contributions to lepton physics, specifically for the discovery of the tau lepton";

and

•Reines, Frederick, U.S.A., University of California at Irvine, Irvine, CA, "for pioneering experimental contributions to lepton physics, specifically for the detection of the neutrino".

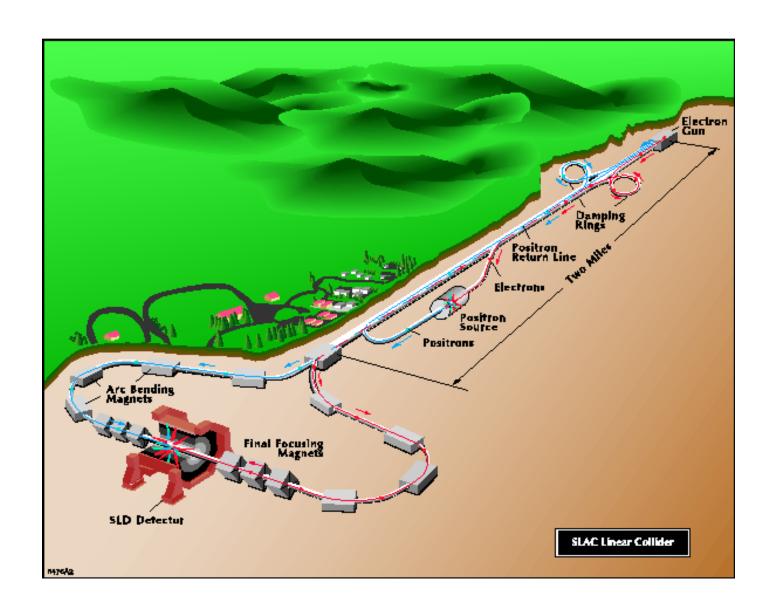
Tau: The 3rd Electron-like Particle

In 1975, Martin Perl (SLAC) scanned the 1973-1974 SPEAR experimental data, searching for a particularly unusual type of **event**. What if, he supposed, sometimes an electron and positron annihilate, and the detector records only one electron-type track, and one muon-type track?

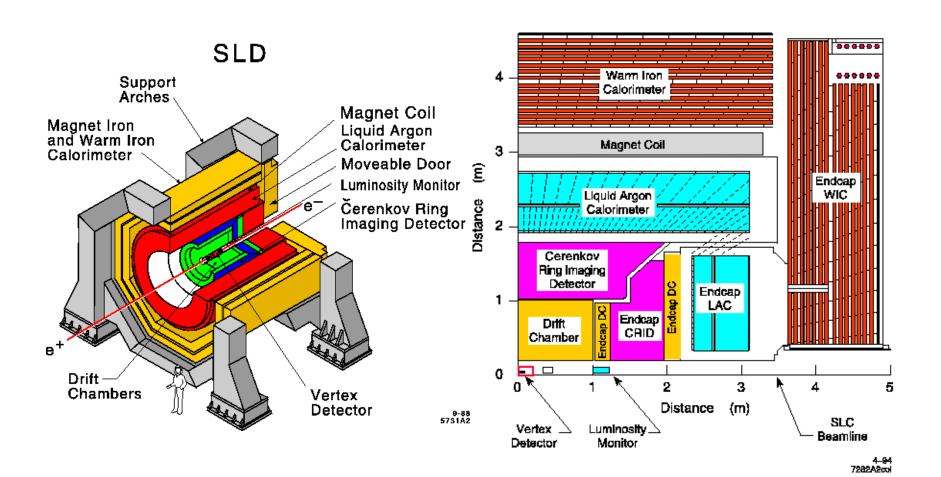
These theoretically-predicted events were found, and at rates that could only be explained by postulating another new particle type, one just like the electron but 3,000 times more massive. (The muon, too, is just like the electron, but 200 times more massive and no one -- yet -- understands why there are three electron-like particles.)



SLD + SLC



SLAC

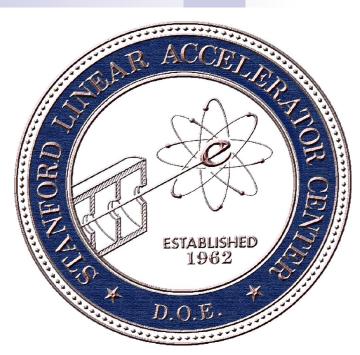


SLAC





Laboratory Report: SLAC





Jonathan Dorfan

ICFA Seminar, Kyungpook National University, Daegu, Korea. September 28th, 2005



SLAC Future – Responding to the Changing Scientific Landscape

The balance of the scientific elements of the Lab are changing:

The size of the photon science program will grow significantly in the next three years. By 2009, the on-site accelerators, SPEAR3 and Linac Coherent Light Source (LCLS), will both be doing Photon Science

B Factory will run through FY2008

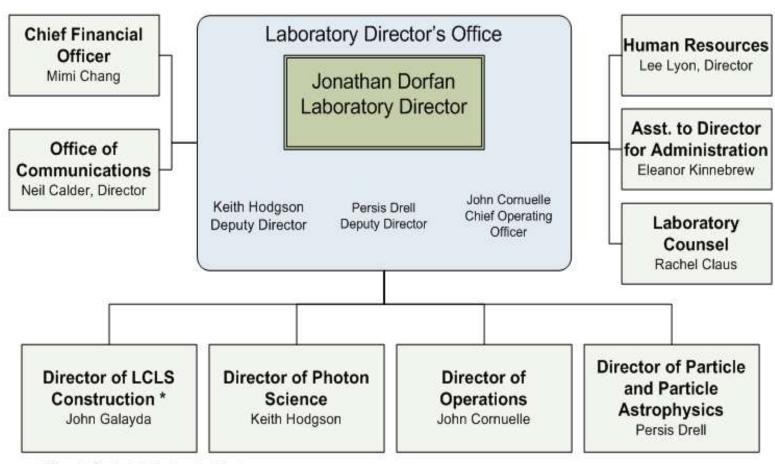
Post FY2008, Particle and Particle Astrophysics will be focused on: BABAR analysis, the ILC machine and detector, accelerator science/R&D, Gamma ray Large Area Space Telescope (GLAST), Large Synoptic Survey Telescope (LSST), Enriched Xenon Observatory and particle/astro/cosmology theory

Scientific Computing challenges abound in both focus areas

As of May 2005, the SLAC Organization has been changed.

Stanford Linear Accelerator Center

Directorate Level Organization



^{*} Reports directly to the Laboratory Director





Mission

Photon Science Discoveries

 To make discoveries in photon science at the frontiers of the ultrasmall and ultrafast in a wide spectrum of physical and life sciences

Particle and Astroparticle Physics Discoveries

 To make discoveries in particle and astroparticle physics to redefine humanity's understanding of what the universe is made of and the forces which control it

Operate Safely; Train the Best

 To operate a safe laboratory that employs and trains the best and brightest, helping to ensure the future economic strength and security of the nation

SLAC as an International Research Facility

- 3000 scientists from ~25 nations use SLAC facilities to do their research
- 1400 Laboratory staff
- Annual budget ~\$250M











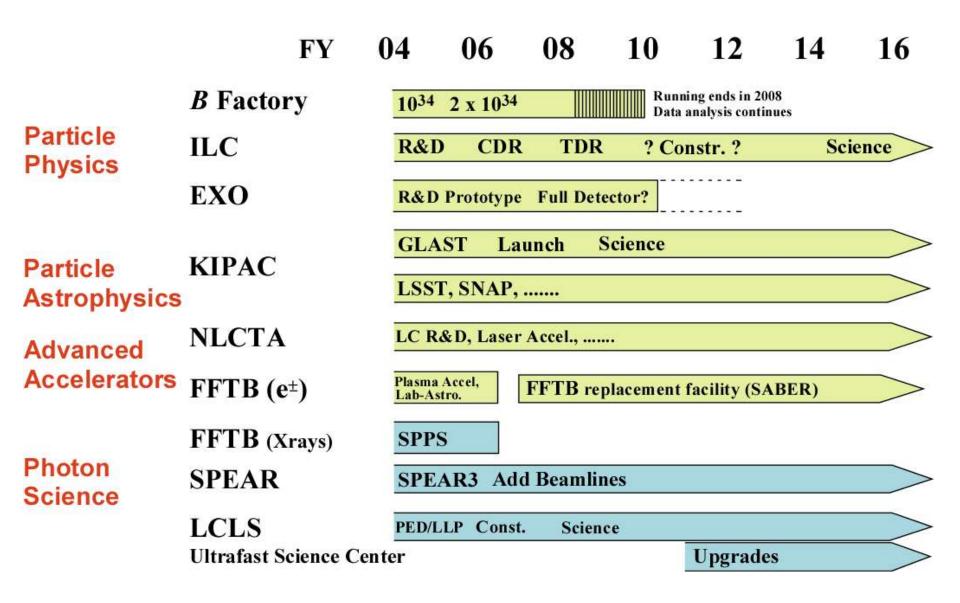




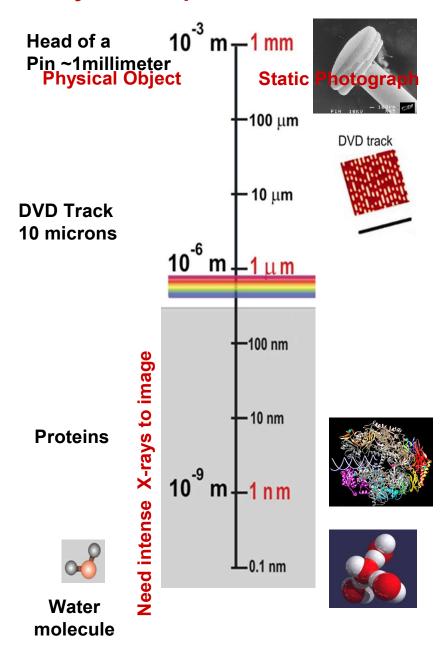




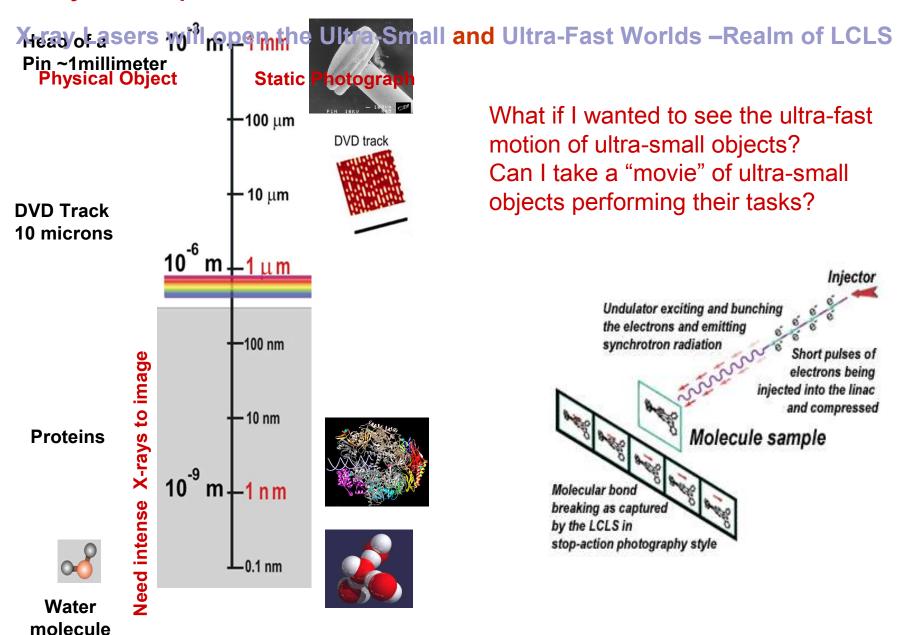
SLAC Future Experimental Program



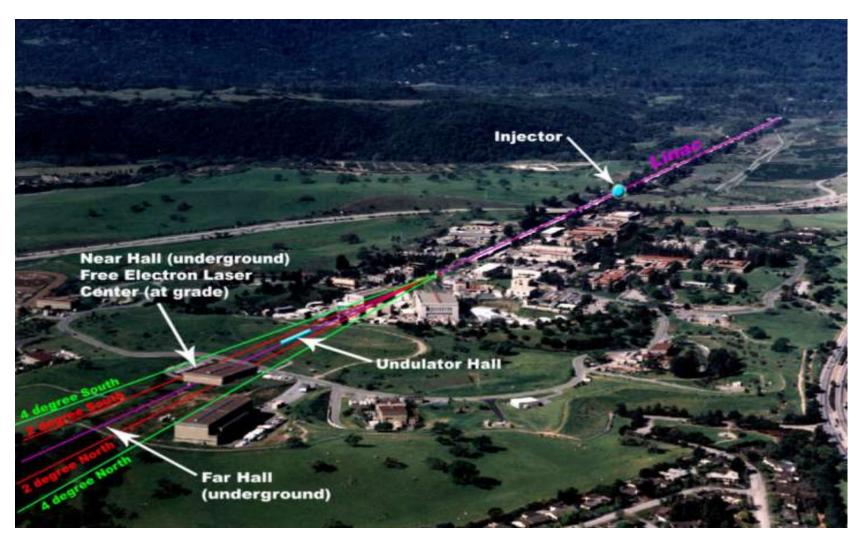
X-Rays have opened the Ultra-Small World -- Realm of SPEAR3



X-Rays have opened the Ultra-Small World -- Realm of SPEAR3



Linac Coherent Light Source at SLAC



LCLS Will Be The World's First X-ray Laser

LCLS Construction

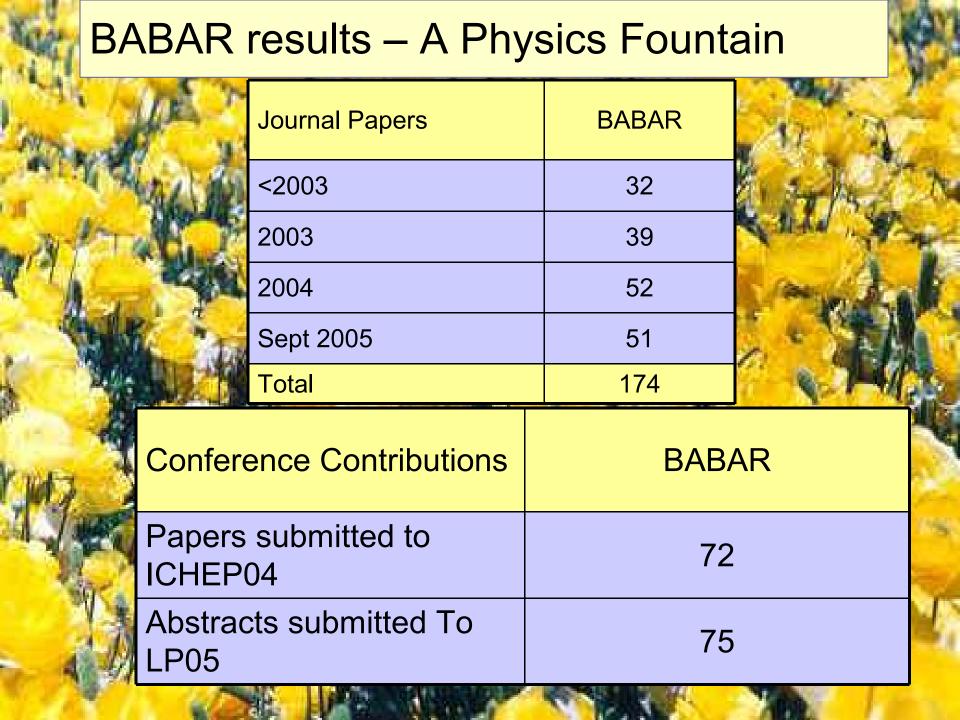
Budget - \$379M through 2009. Commissioning in FY2008.
 Operations start in April 2009

LinacCoherentLightSourceFundingProfile(AYM\$)										
	FY02	FY03	FY04	FY05	FY06	FY07	FY08	FY09	Total	
TEC Funding	0.00	5.93	7.46	50.08	85.54	105.50	5050	10.00	315.00	
OPC Funding	1.50	0.00	2.00	4.00	3.50	16.00	15.50	21.50	64.00	
TotalFunding	1.50	5.93	9.46	54.08	89.04	121.50	66.00	31.50	379.00	

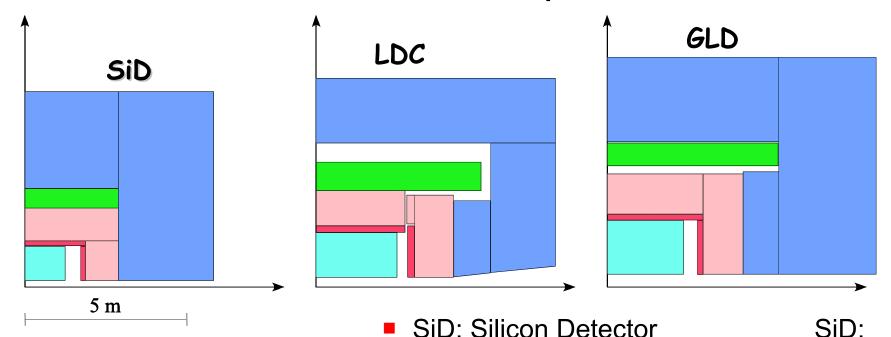
- LCLS is a major endeavor. Will achieve femtosecond time-scales at 10⁹ times the instantaneous brightness of third generation light sources
- Anticipate major upgrades to begin ~2012
- Stanford/SLAC in partnership with DOE has created the Ultrafast Science Center (UFC) to provide a broad-based intellectual focus to underpin the LCLS program.
 - Professor Phil Bucksbaum will move from Michigan to Stanford/SLAC as the first Director of UFC

Scientific Focus of Current and Future SLAC Particle/Astroparticle Program

Current and planned SLAC HEP program is addressing compelling scientific questions facing the field
□ Where did the antimatter go? (B-Factory)
☐ Are there new symmetries and forces of nature? (B-Factory, ILC
□ Why are there so many particles? (B-Factory)
□ What is Dark Matter? (LSST,GLAST,ILC)
□ Can we solve the mystery of Dark Energy? (LSST, JDEM, ILC)
☐ Is there grand unification of particles and forces? (ILC, EXO)
□ What are neutrinos telling us? (EXO)
□ Are there extra dimensions of space? (ILC)
Doing accelerator research and technology development to meet current challenges and for the longer term future of the field
□ PEP-II
□ Multi-TeV LCHigher Gradient + Two Beam acceleration
□ Future acceleration concepts (laser and plasma-wakefield)



ILC Detector Concepts Three + 1 detector concepts



Main Tracker **EM Calorimeter Had Calorimeter** Cryostat / Solenoid Iron Yoke / Muon System

- Small, 'all' silicon
- LDC: Large Detector Concept
 - B R²

BR²

TPC based

SIGLO: habbahtargin Detectorent with DiD B_{R²}

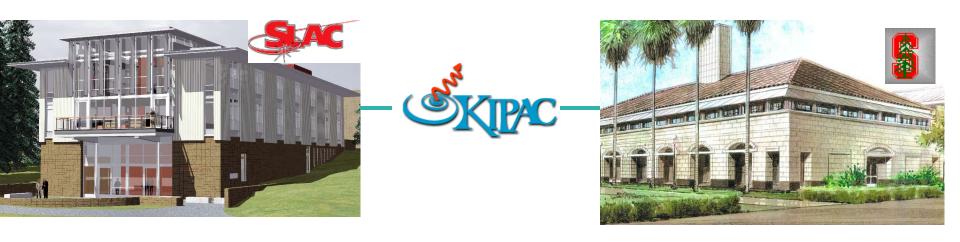
LDC:

v

SLAC People Working on SiD

- Simulation
 - □ N. Graf, T. Johnson, R. Cassell, J. McCormick
- MDI & Backgrounds
 - M. Woods, T. Maruyama, T. Markiewicz, K. Moffeit
- EMCal
 - ☐ G. Haller, D. Freytag, R. Herbst, mb
- Tracker Studies
 - □ T. Nelson, J. Jaros
- Physics Benchmarks
 - □ T. Barklow
- VXD Studies
 - □ Su Dong

Kavli Institute forParticle Astrophysics and Cosmology



Founded 2003

Director: Roger Blandford

Deputy Director: Steve Kahn

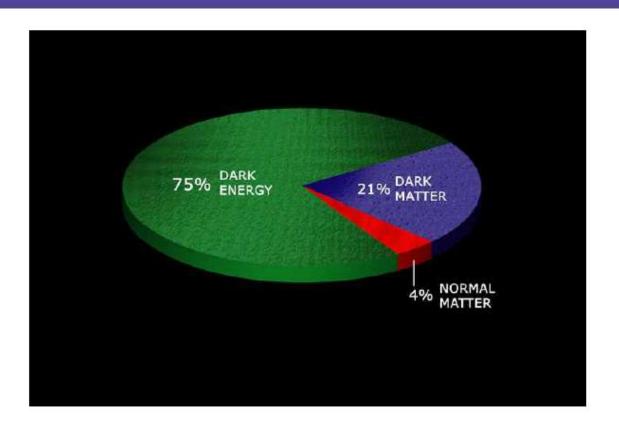
~60 members (30 new)

Two new buildings, labs

Instrumentation, data analysis, particle astrophysics, relativity, computational astrophysics, observational cosmology, theoretical cosmology...

KIPAC is a major commitment by Stanford

What is the Universe made of?



- This discussion is only about the 4 % 'ordinary matter'
- We hopefully are on the verge of finding out what the 21 % dark matter are made of

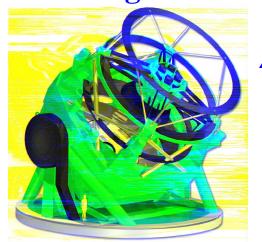
Construction of Kavli Building

Occupancy January 2006



KIPAC Projects

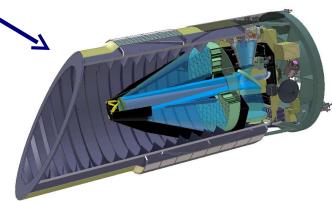
LSST First Light ~ 2012



GLAST Launch 2007 **Dark Energy and Matter**

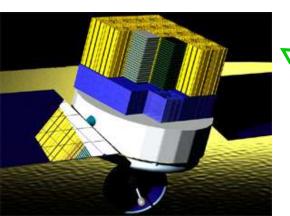
All Sky

High Resolution SNAP Launch ~ 2014



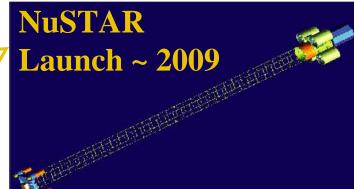
Combining:

- •Physics and Astronomy
- •Theory and Experiment
- •SLAC and Campus
- •DOE, NASA and NSF



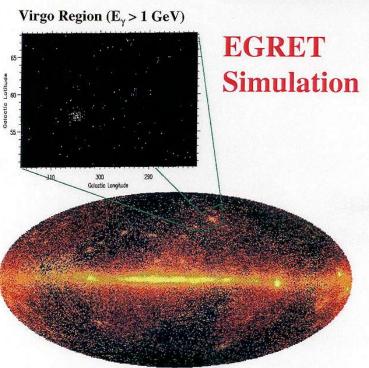
All Sky Resolution

Cosmic Accelerators and Black Holes

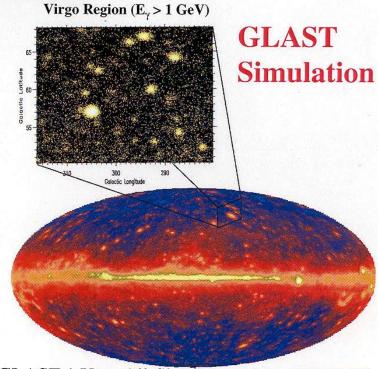


GLAST: Gamma-ray Large Area Space Telescope





EGRET 1 Year All-Sky Survey (E_y> 100 MeV)



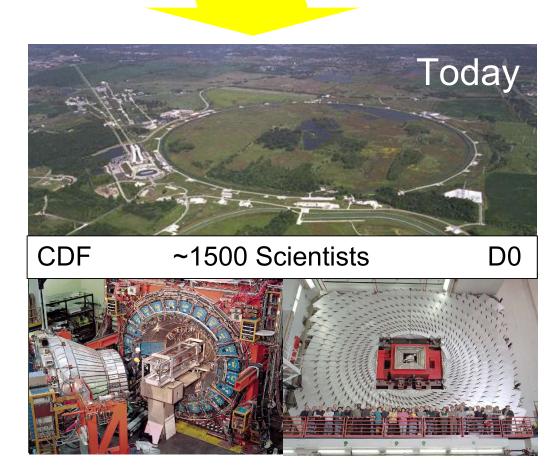
GLAST 1 Year All-Sky Survey (E_y> 100 MeV)

Fermi National Adderator Laboratory

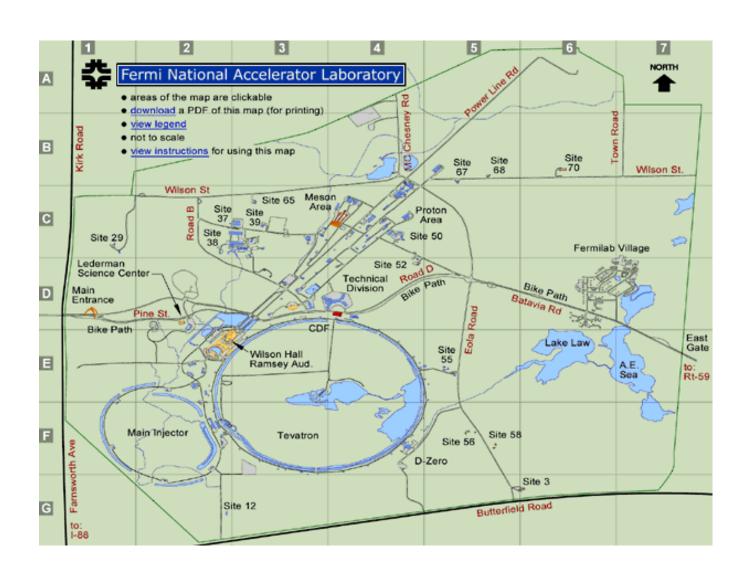


Many generations of Accelerators created with higher and higher energy given to the beam particle



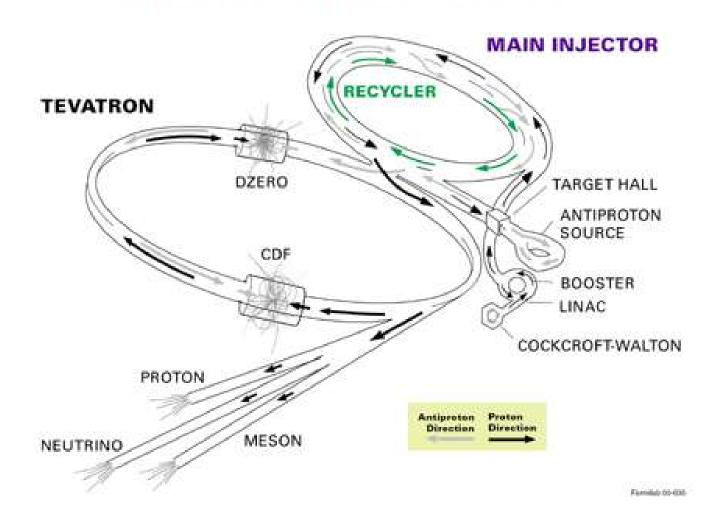


FNAL

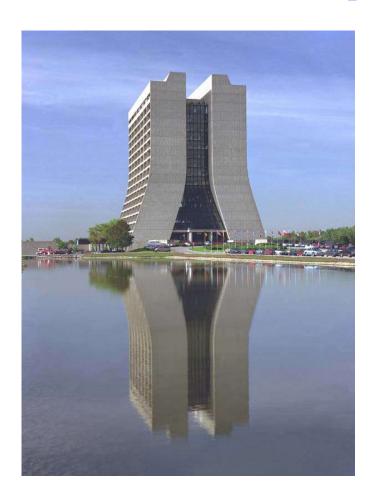


FNAL

FERMILAB'S ACCELERATOR CHAIN



FNAL





FNAL (CDF & D0)





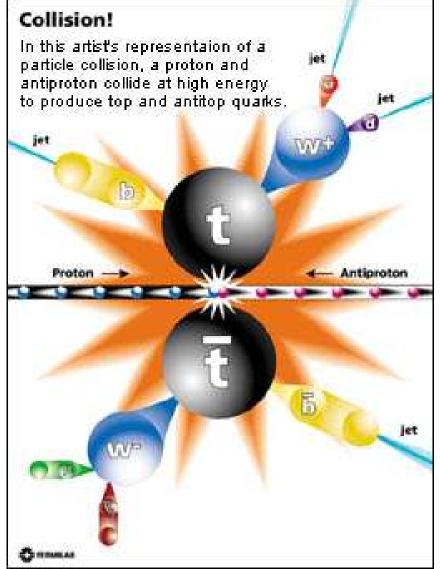
FNALBudgets and Statistics

Type of funding	Budget (\$ millions)	Expenditure (\$ millions)
Operating	240	235
Capital equipment	41	47
Construction/Plant	26	39
Totals	307	321

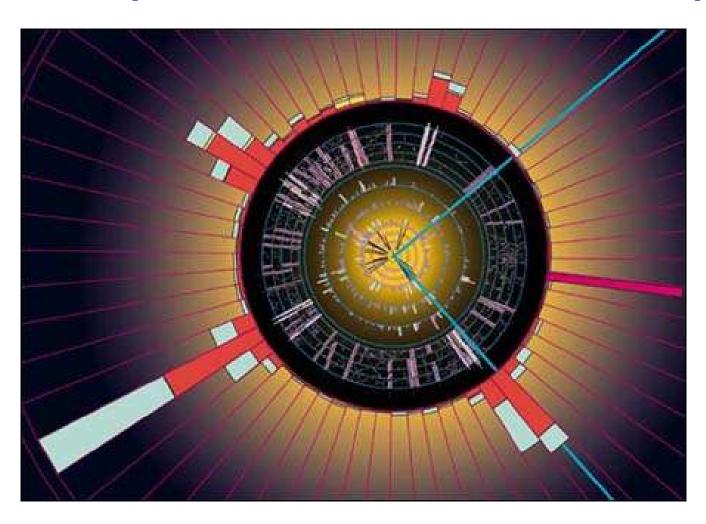
There are approximately 2200 people working at Fermilab, including Cooperative students.

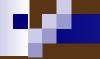
Fermilab has also almost 2300 users. A user is a researcher from an outside institution who uses Fermilab to carry out research in particle physics or related disciplines.

FNAL (T – quark observation)

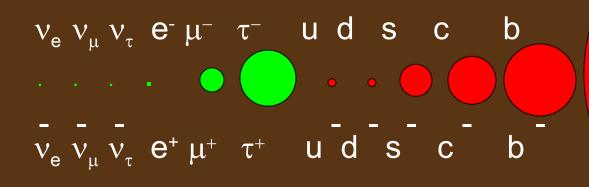


FNAL (T – quark observation)



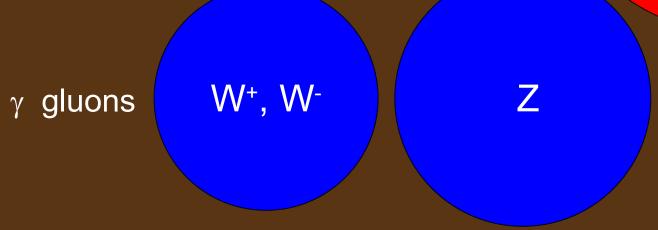


Elementary Particles and Masses



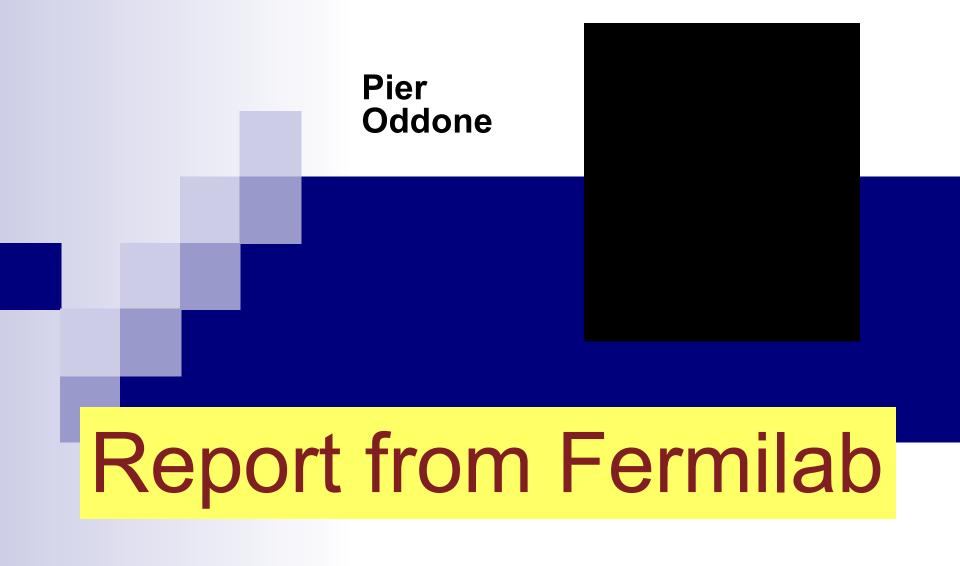
top quark

anti-top quark



(Mass proportional to area shown but all sizes still < 10⁻¹⁹ m)

Why are there so many? Where does mass come from?



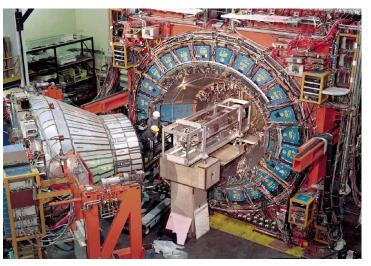
Presentation to ICFA Symposium Daegu, Korea September 2005

The "ships of the line"

- In the near term:
 - □ Tevatron Program CDF and D0
 - ☐ The neutrino program MiniBoone and MINOS
 - □ Large Hadron Collider and CMS
- In the future:
 - □ Neutrinos: NOvA
 - International Linear Collider

Tevatron Program

- Greatest window into new phenomena until LHC is on
- 1500 collaborators, 600 students + postdocs
- Critically dependent on Luminosity
- Doubling time a major consideration





CMS: Compact Muon Detector Output Out

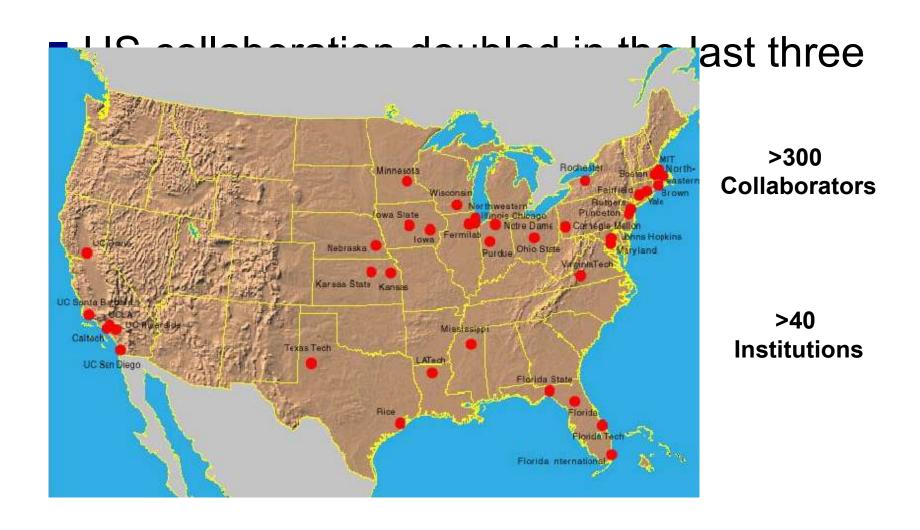


Muon detectors



Magnet cold mass

CMS: Compact Muon Detector

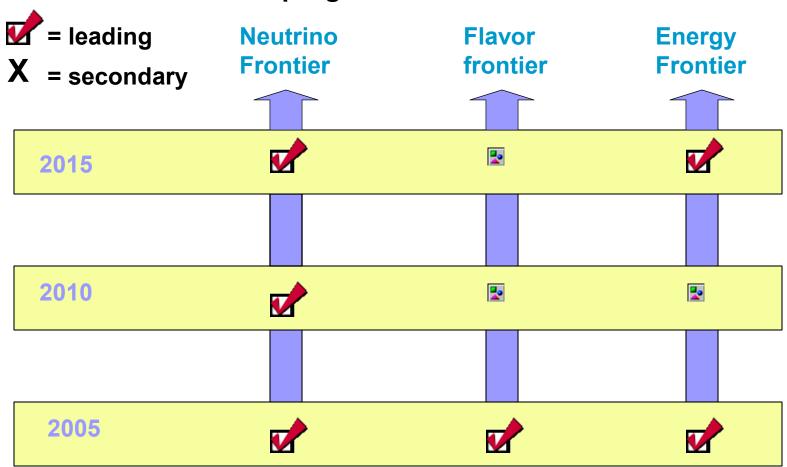


US CMS: and Fermilab's role

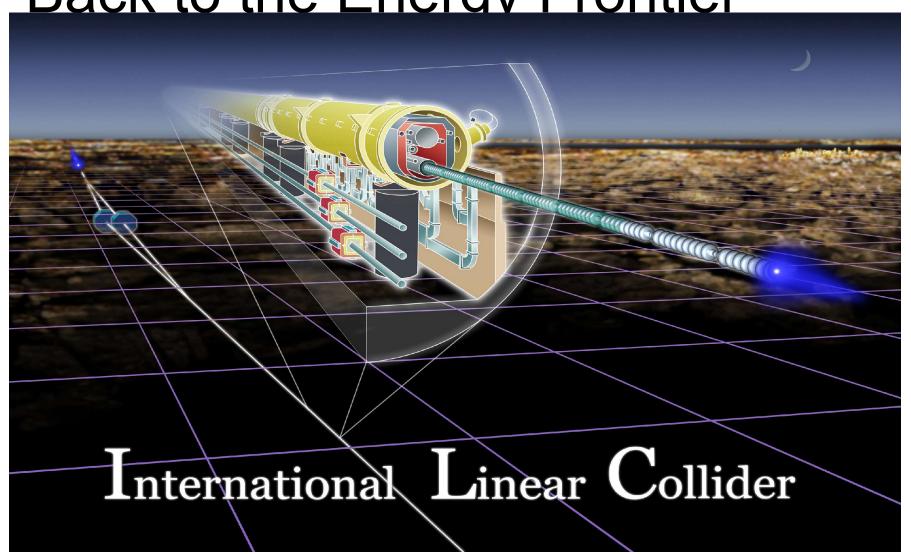
- Only major US lab associated with CMS: a central support role for the US community
- This was the case during construction
- Attention now to huge data and physics discovery challenge: the LHC Physics Center (LPC)

Strategic context: U.S. contribution

Domestic accelerator program with new and redirected investment



Back to the Energy Frontier



The Energy Frontier (ILC)

Goals:

- Establish all technical components, costs, engineering designs, management structures to enable "early" decision (by 2010) as part of the global effort.
- Position US (and Fermilab) to host the ILC.
- Position US (and Fermilab) to play major roles in detector development and physics analysis.
- This is the highest priority initiative for the laboratory, WHEREVER the ILC is finally built